Name: Stud	y Guide
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Topic: Astronomy

The Universe

Biggest to smallest

Universe, Galaxy, Solar System, Sun, Planet, Moon

Origin of the Universe

- o <u>Big Bang</u>: theory that a giant "explosion" some 13.8 billion years ago created the universe, and that it has been expanding from this explosion ever since.
- o Evidence of the Big Bang Theory:
 - Cosmic background radiation: can still be detected (in all directions) from the original big bang
 - **Stellar radiation**: star light is affected by a star's motion.
 - o **Doppler Effect:** shift in light/ sound waves as the source moves either toward or away from the observer.

<u>Red Shift:</u> waves are spread out when object is moving <u>away</u> from the observer

<u>Blue Shift:</u> waves are compressed when object is moving <u>toward</u> the observer

Spectra of galaxies shows a shift toward the <u>Red end</u> of the spectrum due to the Doppler Effect.

This means most galaxies are moving **AWAY** from Earth.

Blue		Red
Laboratory Hydrogen Spectral Lines		
Blue		Red
Galaxy A Spectral Lines		
Blue		Red
Galaxy B Spectral Lines		
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Stars

- o **Galaxies:** billions of stars
 - Shapes include spiral, elliptical, and irregular
 - There are billions of galaxies in the universe

Classifying Stars

- Stars are grouped by comparing their **luminosity** to their **temperature**. Check out page 15 in the ESRT
- o Star Types:
 - Main Sequence Stars: the diagonal band across the star graph where most plotted stars are found
 - **Giant Stars**: bigger, brighter stars that are older than the main sequence stars
 - **Super Giants:** 100 to 1000 time the diameter of the sun
 - White Dwarfs: old stars that have collapsed, making them very small, very dense, and very hot

Life Cycle of a Star

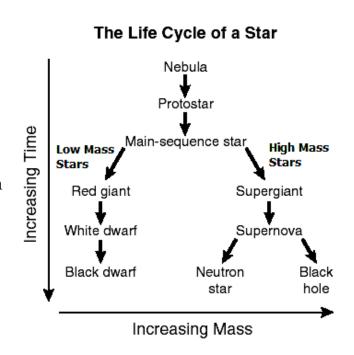
- 1. Stars originate from clouds of dust and gas molecules, known as a **nebula**.
- 2. Gravity caused the dust and gas to clump together forming a **protostar**.
- 3. Inside the core:
 - Temperature, gas pressure, and density increases as atom collisions increase.
 - Nuclear fusion starts!
 - o <u>Nuclear fusion:</u> the process of stars making energy by combining hydrogen atoms to make helium.

HAPPY LIFE

- 1. Equilibrium between gravity and heat and pressure from fusion must be kept in order for a star to live peacefully.
- 2. A star will live happily as long as it continues to burn hydrogen.

DEATH OF A STAR

- 1. Gravity, heat and pressure are no longer in equilibrium.
- 2. What happens after this point depends on the mass of the original star. Look at the chart →



Solar System

- o <u>Solar System:</u> the sun and all objects that orbit the sun under its gravitational influence.
- The **Solar System formed about 4.6 billion years ago** from dust/ gas cloud pulled together by gravity.

Gravity: influences the motion of celestial objects

- Gravity is directly proportional to an object's mass.
 - o The higher the mass of a planet, the more gravitational pull it has
- Gravity is indirectly proportional to an object's distance
 - The farther two objects are away from each other, the less gravity between them.

Asteroids, comets, and meteors

- o **Asteroids:** a solid rock/ metallic bodies that orbit around the sun.
 - Found between Mars and Jupiter.
- o Comets: objects composed of rock and ice that easily vaporize
 - Comets move around the sun.
- o <u>Meteors:</u> path of light from meteoroid burning as it passes through Earth's atmosphere; also called a shooting star.
 - Impact events have been correlated with mass extinction and global climate change.